

Direct and/or inverse energy cascades in TURBULENCE

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HOHAI UNIVERSITY (MARCH 2018)



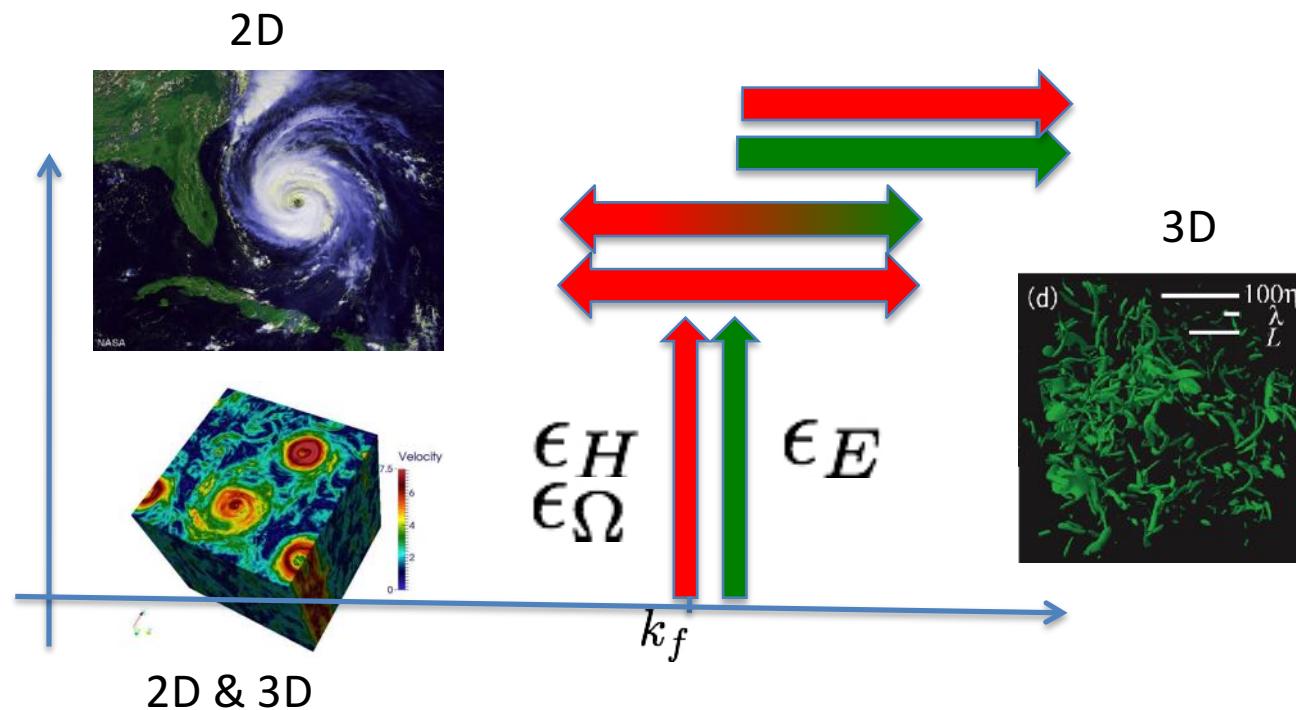
Credits [in order of appearance]: **F. Toschi** (TuE, The Netherlands); **S. Musacchio** (CNRS-Nice, France); **E. Titi** (Weizmann Institute of Science, Israel), **G. Sahoo** (University of Helsinki, Finland); **F. Bonaccorso, M. Buzzicotti** (Univ. of Roma ‘Tor Vergata’, Italy), **M. Linkmann** (Univ. of Marburg, Germany); **A. Alexakis** (ENS-Paris, France)

MOTIVATIONS:

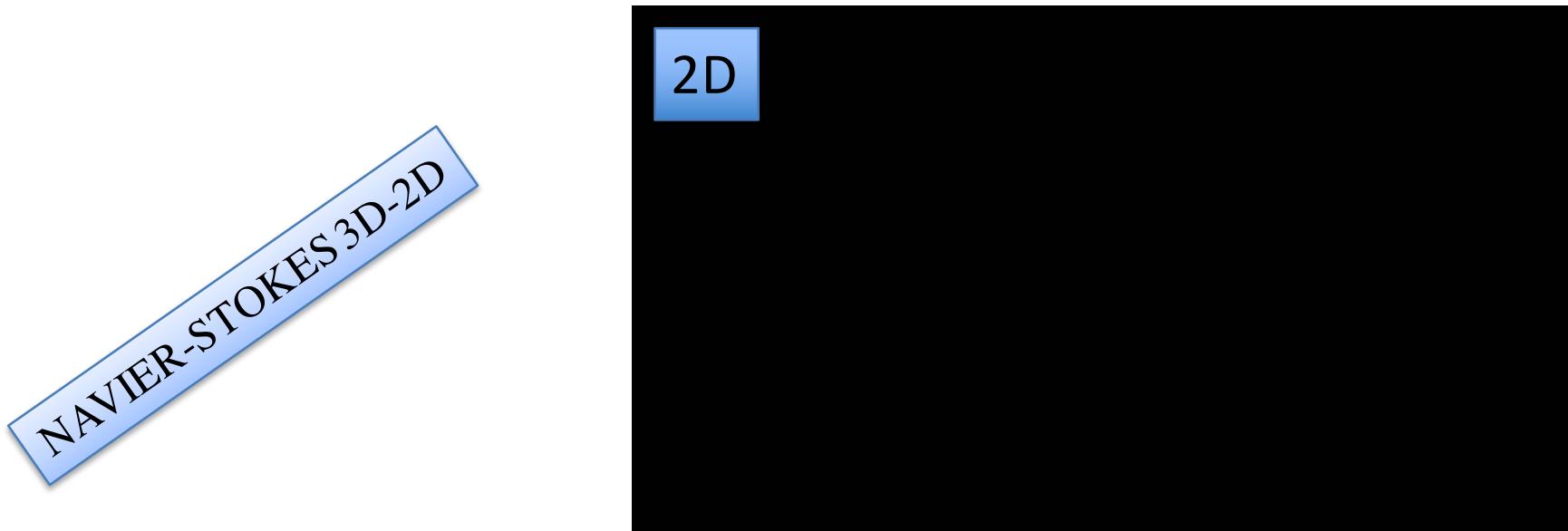
Q1: HOW TO PREDICT THE DIRECTION OF THE TRANSFER (FORWARD/BACKWARD) AND ITS ROBUSTNESS UNDER EXTERNAL PERTURBATION (FORCING/BOUNDARY CONDITIONS)?

Q2: WHAT ABOUT FLUCTUATIONS AROUND THE MEAN?

AS A MATTER OF FACT, FOR 3D NAVIER STOKES EQUATIONS, WE DO NOT KNOW HOW TO PREDICT NEITHER THE SIGN OF THE MEAN ENERGY TRANSFER NOR THE INTENSITY OF THE FLUCTUATIONS AROUND IT.



(NASA/Goddard Space Flight Center Scientific Visualization Studio)



NAVIER-STOKES3D-2D

Entry #: 84174

3D

Vortices within vortices:
hierarchical nature of vortex tubes in turbulence

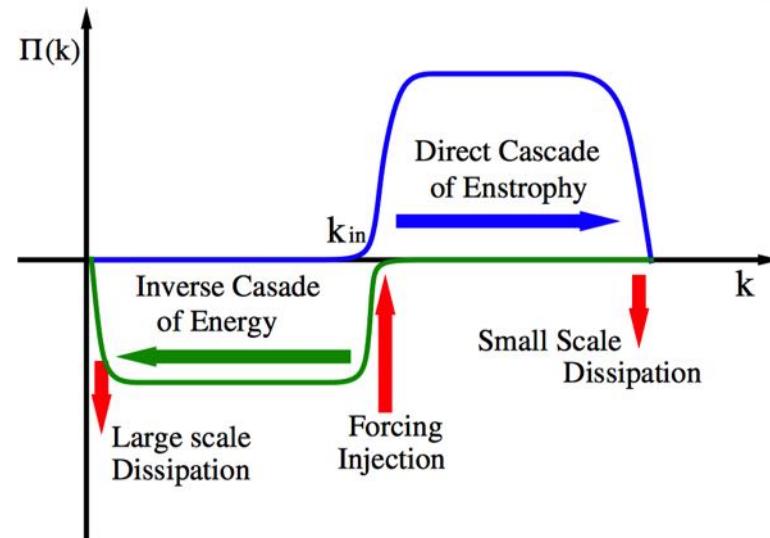
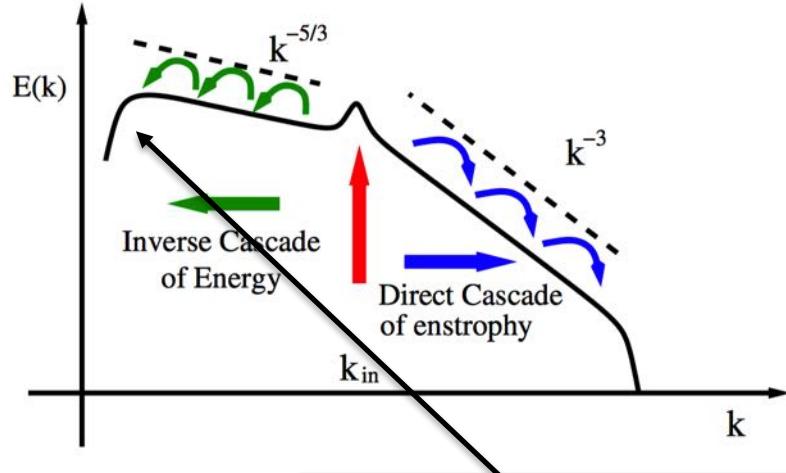
Kai Bürger¹, Marc Treib¹, Rüdiger Westermann¹,
Suzanne Werner², Cristian C Lalescu³,
Alexander Szalay², Charles Meneveau⁴, Gregory L Eyink^{2,3,4}

¹ Informatik 15 (Computer Graphik & Visualisierung), Technische Universität München
² Department of Physics & Astronomy, The Johns Hopkins University
³ Department of Applied Mathematics & Statistics, The Johns Hopkins University
⁴ Department of Mechanical Engineering, The Johns Hopkins University

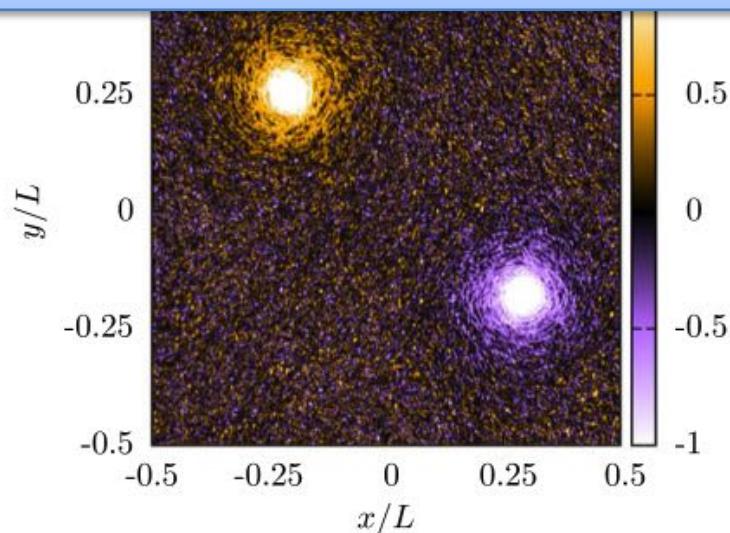
2 invariants (positive definite)

2D

$$\Omega = \int d^3x \boldsymbol{\omega} \cdot \boldsymbol{\omega} \quad E = \int d^3x \mathbf{v} \cdot \mathbf{v}$$



INVERSE CASCADE + IR CUT-OFF →
→ LARGE SCALE CONDENSATE

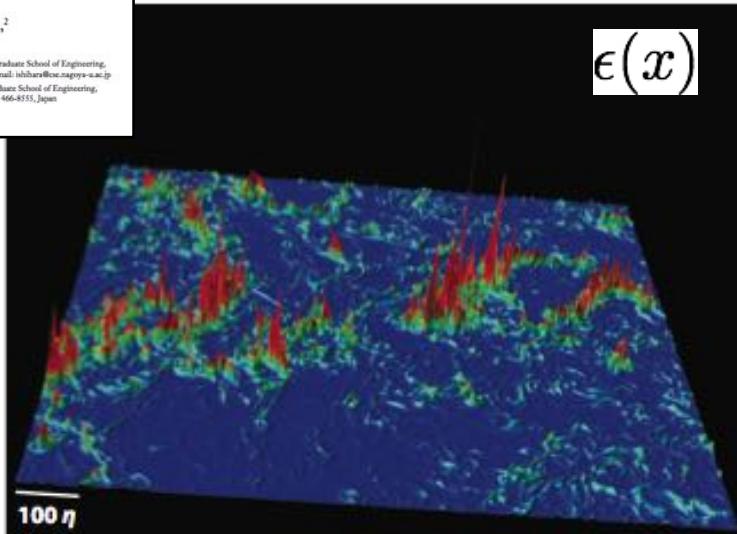


Study of High-Reynolds
Number Isotropic Turbulence
by Direct Numerical
Simulation

Takashi Ishihara,¹ Toshiyuki Gotoh,²
and Yukio Kaneda¹

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3D HOMOGENEOUS AND ISOTROPIC TURBULENCE FLUCTUATIONS: SMALL-SCALES INTERMITTENCY



3D

DIRECT
ENERGY
CASCADE

$$E = \int d^3x \mathbf{v} \cdot \mathbf{v}$$

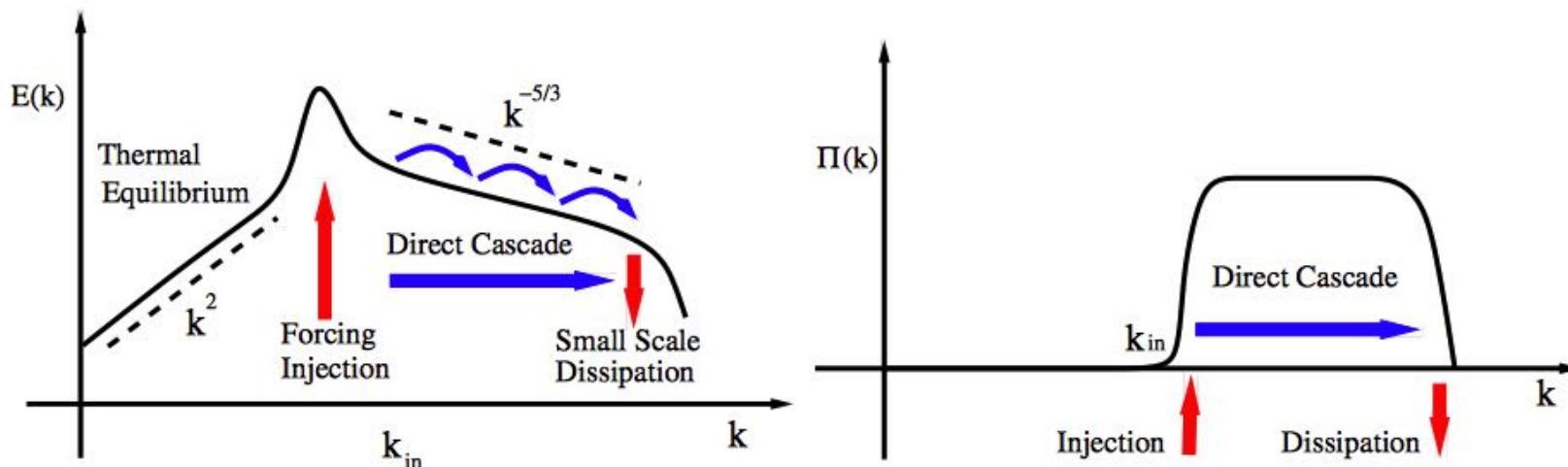
$$H = \int d^3x \boldsymbol{\omega} \cdot \mathbf{v}$$

HELICITY

2 invariants (only one positive definite)

Figure 4

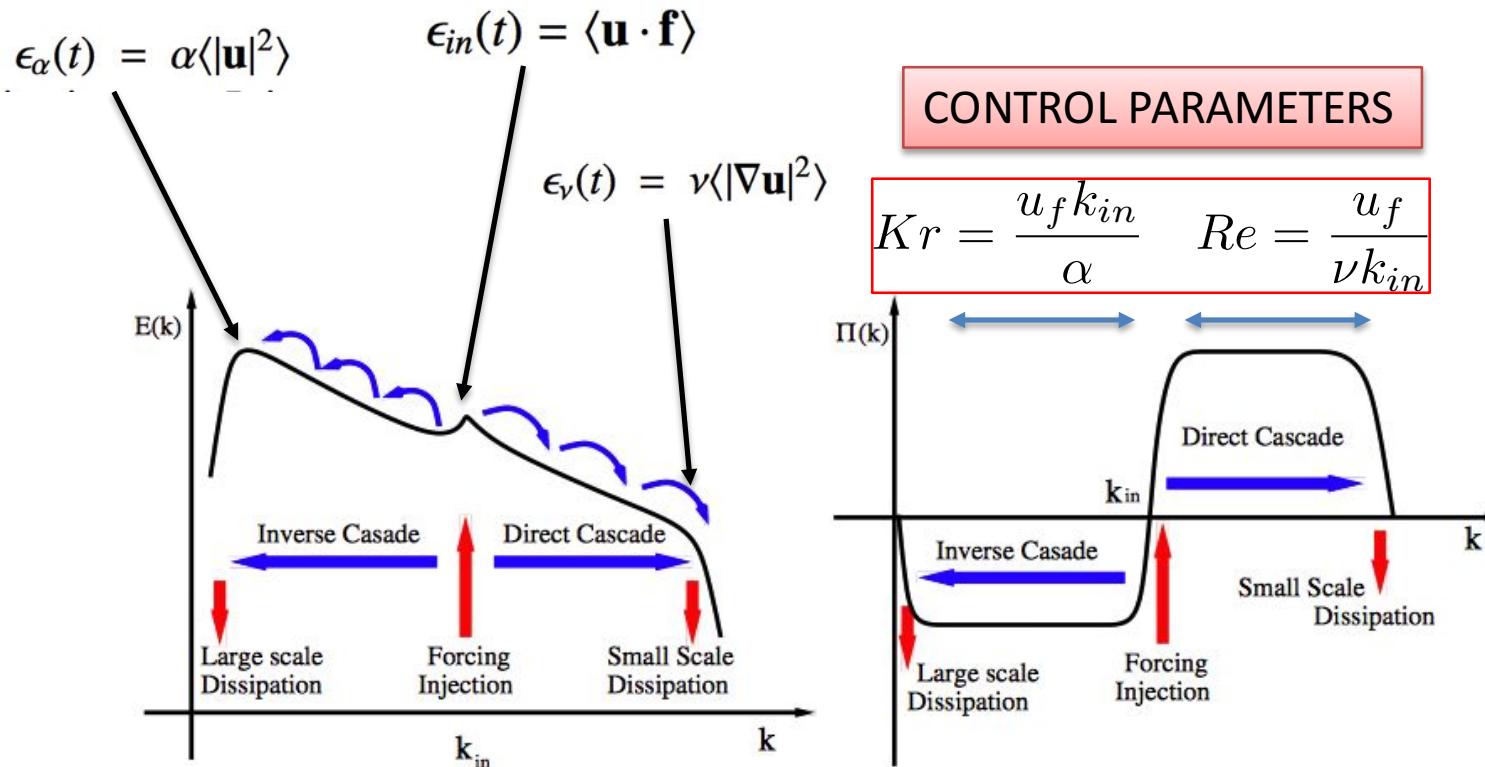
Snapshot of the intensity distributions of (a) the energy-dissipation rate $\tilde{\epsilon} = \epsilon/(2v)$ and (b) the enstrophy $\Omega = \omega^2/2$ on a cross section in DNS-ES at $R_\lambda = 675$ in arbitrary units.



2D/3D NOT THE END OF THE HISTORY: SPLIT ENERGY CASCADE

$$\partial_t \mathbf{u} + \mathbf{u} \cdot \nabla \mathbf{u} = -\nabla P + \nu \Delta \mathbf{u} - \alpha \mathbf{u} + \mathbf{f}.$$

$$\partial_t \mathcal{E} = \epsilon_{in}(t) - \epsilon_\nu(t) - \epsilon_\alpha(t)$$



DIRECT/INVERSE CASCADE TRANSITION

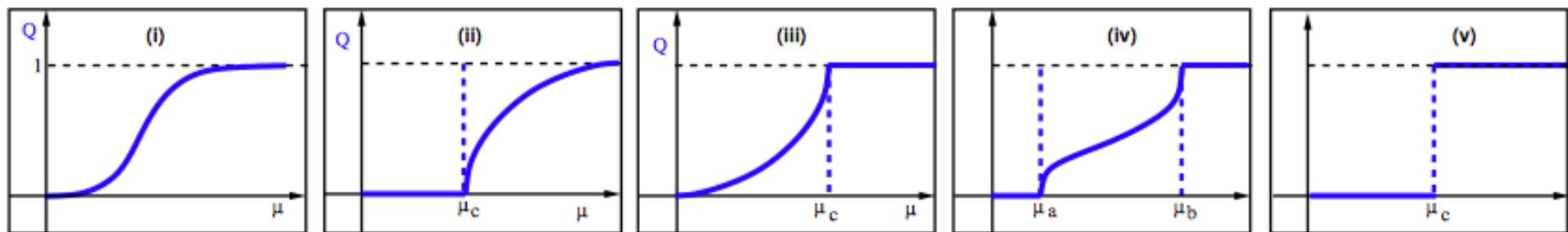
$$\partial_t \mathbf{u} + \mathbf{u} \cdot \nabla \mathbf{u} = -\nabla P + \nu \Delta \mathbf{u} - \alpha \mathbf{u} + \mathbf{f}.$$

$$\partial_t \mathcal{E} = \epsilon_{in}(t) - \epsilon_\nu(t) - \epsilon_\alpha(t)$$

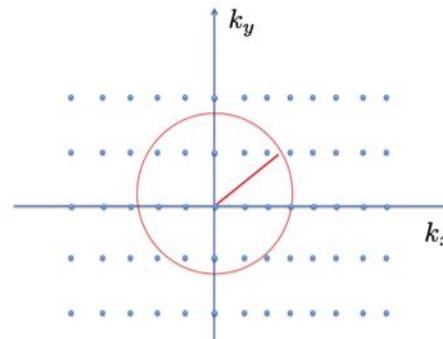
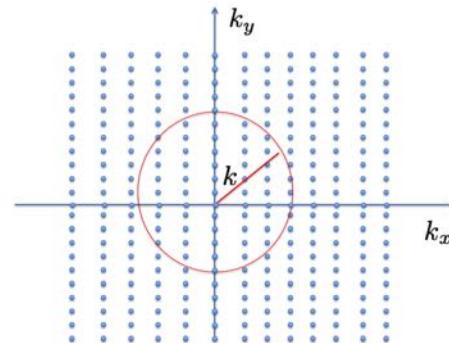
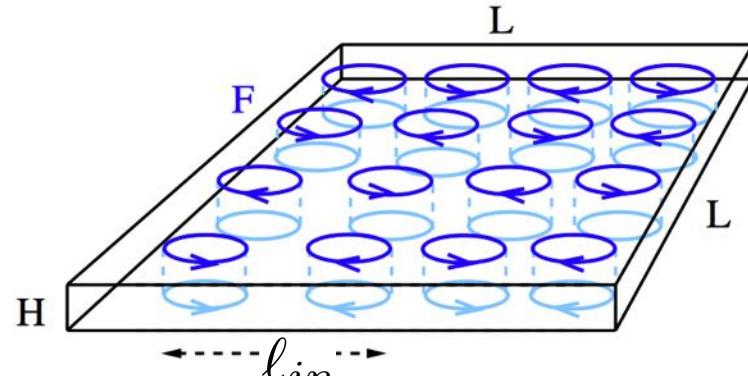
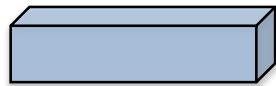
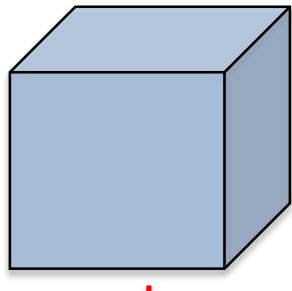
$$\bar{\epsilon}_{in} = \bar{\epsilon}_\nu + \bar{\epsilon}_\alpha \leftarrow \text{STATIONARY BALANCE}$$

$$D_\nu = \frac{\epsilon_\nu}{\mathcal{E}_{in}^{3/2} k_{in}} \quad \text{and} \quad D_\alpha = \frac{\epsilon_\alpha}{\mathcal{E}_{in}^{3/2} k_{in}} \leftarrow \text{DIMENSIONLESS FLUX}$$

$$\lim_{\substack{Re \rightarrow \infty, \\ Kr \rightarrow \infty}} \quad \lim_{\substack{L \rightarrow \infty, \\ k_{max} \rightarrow \infty}} \quad \frac{\epsilon_\nu}{\epsilon_{in}} = Q_\nu > 0,$$



3D $\leftarrow \rightarrow$ 2D: THIN LAYERS



CONTROL PARAMETER

$$\frac{H}{l_{in}}$$

$$\partial_t \mathbf{u}_{2D} + \mathbf{u}_{2D} \cdot \nabla \mathbf{u}_{2D} = -\overline{\mathbf{u}_{3D} \cdot \nabla \mathbf{u}_{3D}} - \overline{\nabla P} + \nu \Delta \mathbf{u}_{2D} - \alpha \mathbf{u}_{2D} + \mathbf{f}_{2D},$$

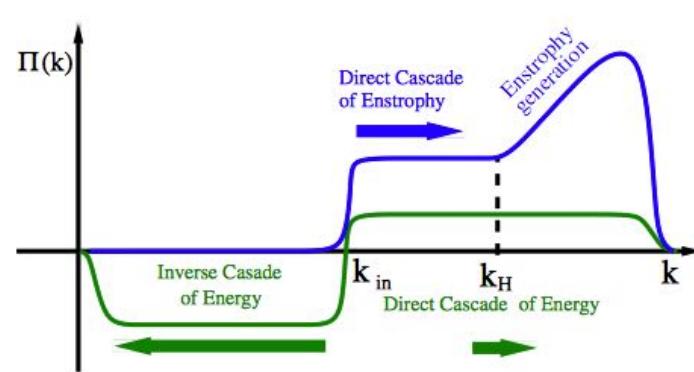
$$\partial_t \mathbf{u}_{3D} + \mathbf{u}_{2D} \cdot \nabla \mathbf{u}_{3D} = -\mathbf{u}_{3D} \cdot \nabla \mathbf{u}_{2D} + (\overline{\mathbf{u}_{3D} \cdot \nabla \mathbf{u}_{3D}} - \mathbf{u}_{3D} \cdot \nabla \mathbf{u}_{3D}) - \nabla P + \nu \Delta \mathbf{u}_{3D} + \mathbf{f}_{3D},$$

ENERGY BALANCE

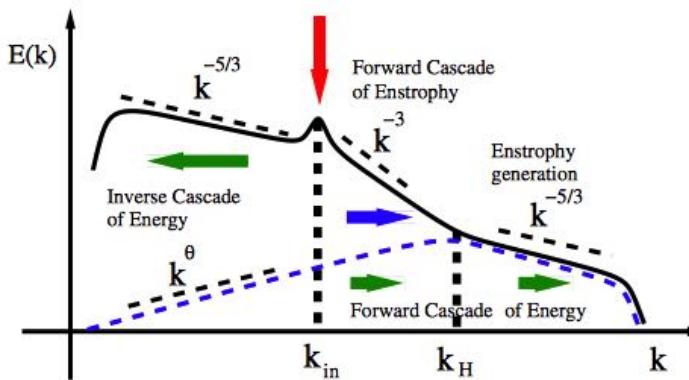
$$\partial_t \mathcal{E}_{2D}(t) = \epsilon_{in}^{2D} - \mathcal{T} - \epsilon_v^{2D} - \epsilon_\alpha^{2D}$$

$$\partial_t \mathcal{E}_{3D}(t) = \epsilon_{in}^{3D} + \mathcal{T} - \epsilon_v^{3D} - \epsilon_\alpha^{3D}$$

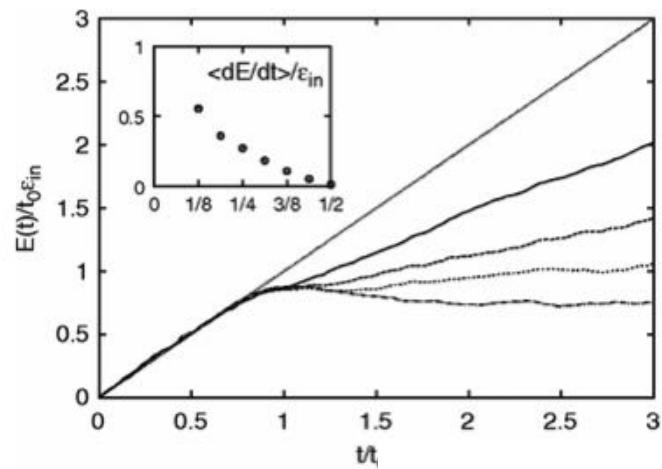
2D < ----- > 3D



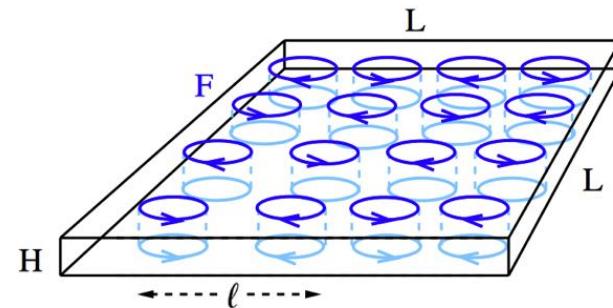
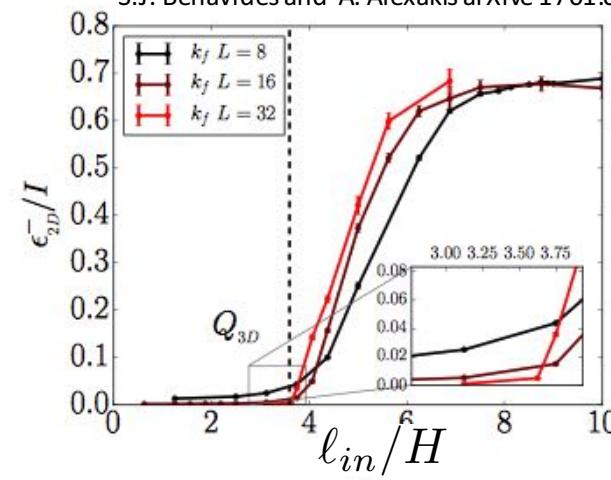
2D < ----- > 3D



A. Celani et al PRL 104, 184605 (2010)



S.J. Benavides and A. Alexakis arXiv 1701.05162



$$\lim H \rightarrow H_c$$

