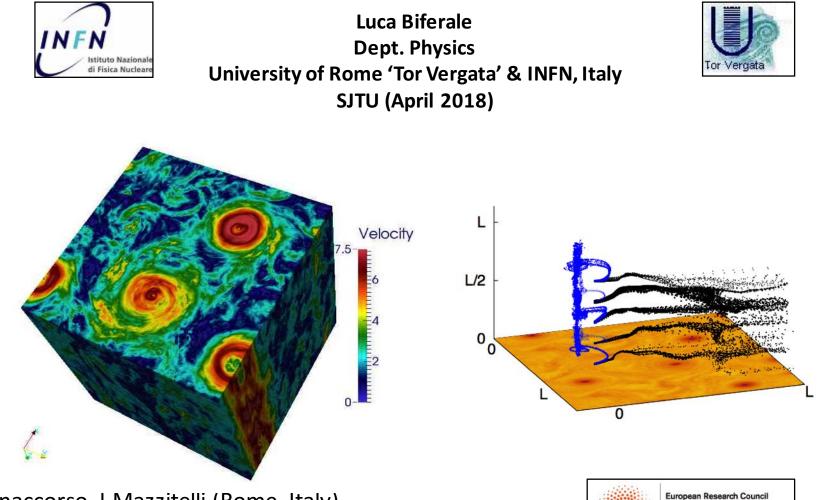
TURBULENT AT HIGH AND LOW ROTATION RATES: EULERIAN AND LAGRANGIAN STATISTICS



F.Bonaccorso, I.Mazzitelli (Rome, Italy) M.Hinsberg, F. Toschi (Eindhoven, The Netherlands) A.Lanotte (Lecce, Italy) S. Musacchio (Nice, France) P.Perlekar (Hydebarad, India)



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PRACE 09_2256 ROTATING TURBULENCE 2015 – 55MH

upporting top researchers rom anywhere in the world

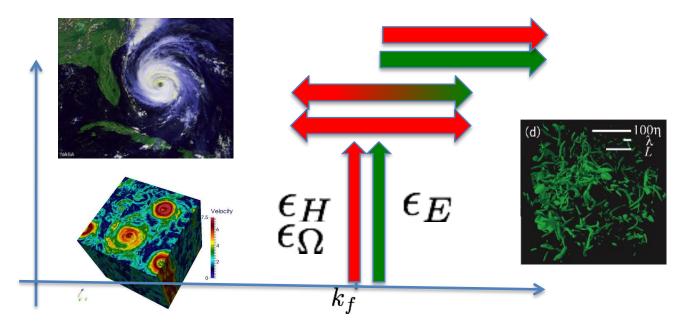
MOTIVATIONS:

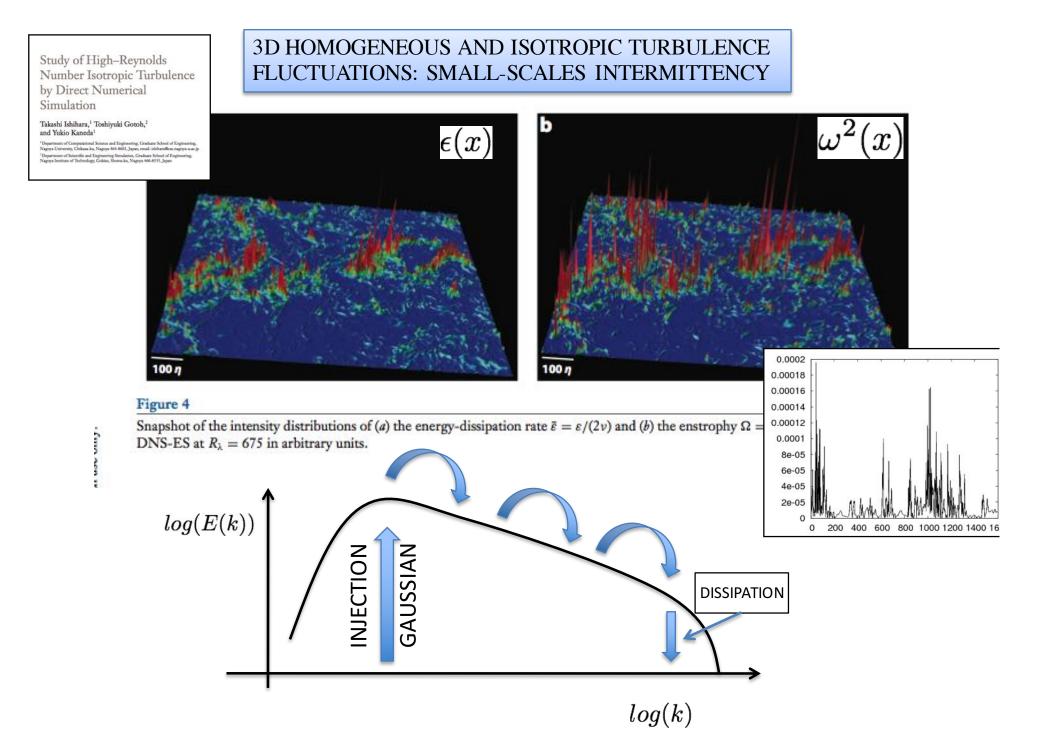
A TALE ABOUT TRANSFER PROPERTIES OF INVISCID CONSERVED QUANTITIES, KINETIC ENERGY, HELICITY ENSTROPHY, MAGNETIC HELICITY ETC...

Q1: HOW TO PREDICT THE DIRECTION OF THE TRANSFER (FORWARD/BACKWARD) AND ITS ROBUSTNESS UNDER EXTERNAL PERTURBATION (FORCING/BOUNDARY CONDITIONS)?

Q2: HOW MUCH THE FLUCTUATIONS AROUND THE MEAN TRANSFER ARE INTENSE AND SELF-SIMILAR (INTERMITTENCY AND ANOMALOUS SCALING) ?

AS A MATTER OF FACT, FOR 3D NAVIER STOKES EQUATIONS, WE DO NOT KNOW HOW TO PREDICT NEITHER THE SIGN OF THE MEAN ENERGY TRANSFER NOR THE INTENSITY OF THE FLUCTUATIONS AROUND IT.





- MOTIVATION: WHY ROTATING TURBULENT FLOWS ARE IMPORTANT

- DIRECT AND INVERSE ENERGY TRANSFERS (2D-3D PHYSICS)

- OUR DNS (DIFFERENCES WRT PREVIOUS STUDIES)

- EULERIAN STATISTICS (MEAN SPECTRAL PROPERTIES)

- EULERIAN STATISTICS (LARGE FLUCTUATIONS)

- LAGRANGIAN STATISTICS (EFFECTS OF CORIOLIS AND CENTRIFUGAL FORCES)

- LAGRANGIAN STATISTICS (SINGLE PARTICLE DISPERSIONS)

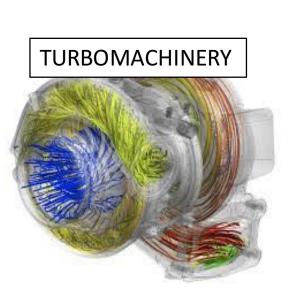
- CONCLUSIONS

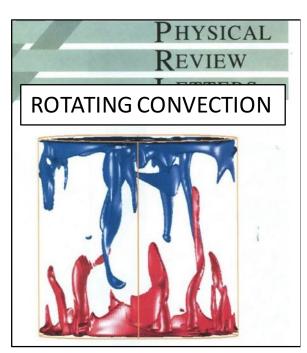
TAYLOR-COUETTE

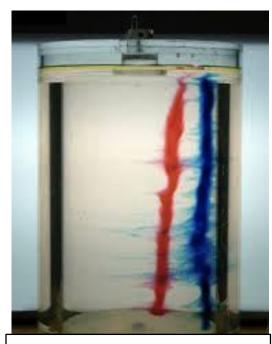
ROTATING CONVECTION (+ STRATIFICATION + MHD)

ROTATING RAYLEIGH-TAYLOR

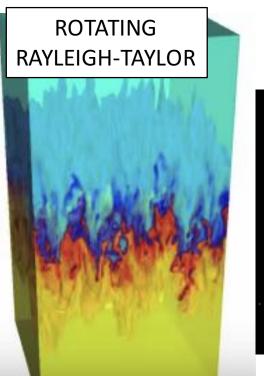
Recent reviews/books by Lohse, Boffetta, Cambon, Clercx, Davidson etc...







CYCLONIC-ANTICYCLONIC DYN.





INNER/OUTER PLANETARY DYNAMICS

NAVIER_STOKES EQS IN A ROTATING FRAME (NO BOUNDARIES)

DNS: A. Pouquet, P. Mininni, A. Alexakis, S. Chen, G. Eyink

$$\partial_t \mathbf{v} + \mathbf{v} \cdot \nabla \mathbf{v} + \mathbf{\Omega} \times \mathbf{v} = \nabla P + \nu \nabla^2 \mathbf{v} + \mathbf{F} - \alpha \mathbf{v}$$

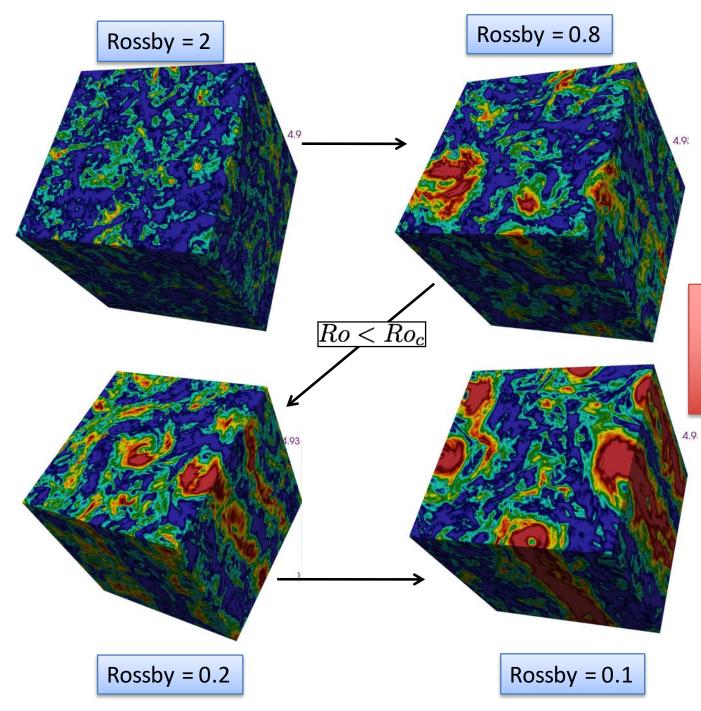
$$oldsymbol{\Omega}$$
 =rotation $P=P_0+rac{1}{2}|oldsymbol{\Omega} imes {f r}|^2$

F-large scale Forcing $\alpha =$ large scale energy sink

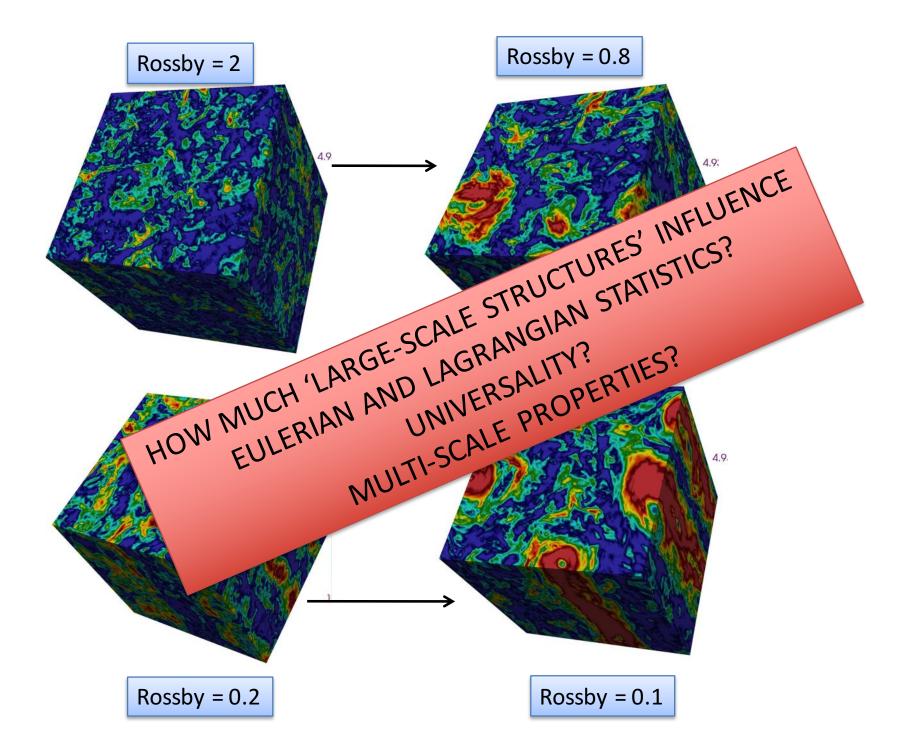
ROSSBY NUMBER ~ NON-LINEAR/ROTATION

$$\operatorname{Ro} \sim \frac{v_0}{\Omega L_0}$$

 $\operatorname{Ro} \geq Ro_c \rightarrow$ forward energy transfer
 $\operatorname{Ro} \leq Ro_c \rightarrow$ forward & backward energy transfer



HOMOGENEOUS ANISOTROPIC 2D & 3D PHYSICS CHOERENT -STRUCTURES



OUR DNS DATA-BASE (EULERIAN + LAGRANGIAN)

NEW FEATURES:

- 1) IDEAL FORCING MECHANISM (AS NEUTRALAS POSSIBLE: ISOTROPIC; NON HELICAL, TIME-COLORED) + LARGE SCALE FRICTION
- 2) UNPRECEDENTED NUMERICAL RESOLUTION/SCALE SEPARATION (UP TO 4096^3)
- 3) LAGRANGIAN STATISTICS (MILLIONS OF TRACERS AND INERTIAL PARTICLES)

| N | Ω | ν | ϵ | ϵ_{f} | u_0 | η/dx | $	au_\eta/dt$ | Re_{λ} | Ro | f_0 | $	au_{f}$ | T_0 | α |
|------|----|---------------------|------------|----------------|-------|-----------|---------------|----------------|------|-------|-----------|-------|----------|
| 1024 | 4 | $7	imes 10^{-4}$ | 1.2 | 1.2 | 1.05 | 0.67 | 120 | 150 | 0.78 | 0.02 | 0.023 | 0.17 | 0.0 |
| 1024 | 10 | $6	imes 10^{-4}$ | 0.46 | 0.59 | 1.6 | 0.76 | 294 | 580 | 0.24 | 0.02 | 0.023 | 0.25 | 0.1 |
| 2048 | 4 | $2.8	imes10^{-4}$ | 1.2 | 1.2 | 1.05 | 0.67 | 380 | 230 | 0.76 | 0.02 | 0.023 | 0.17 | 0.0 |
| 2048 | 10 | $2.2 	imes 10^{-4}$ | 0.45 | 0.64 | 1.7 | 0.72 | 550 | 1170 | 0.25 | 0.02 | 0.023 | 0.3 | 0.1 |
| 4096 | 10 | 1×10^{-4} | 0.46 | 0.65 | 1.7 | 0.78 | 1010 | 1600 | 0.25 | 0.02 | 0.023 | 0.3 | 0.1 |

TABLE I: Eulerian dynamics parameters. N: number of collocation points per spatial direction; Ω : rotation rate; ν : kinematic viscosity; $\epsilon = \nu \int d^3x \sum_{ij} (\nabla_i u_j)^2$: viscous energy dissipation; $\epsilon_f = \int d^3x \sum_i f_i u_i$: energy injection; $u_0 = 1/3 \int d^3x \sum_i u_i^2$: mean kinetic energy; $\eta = (\nu^3/\epsilon)^{1/4}$: Kolmogorov dissipative scale; $dx = L_0/N$: numerical grid spacing; $L_0 = 2\pi$: box size; $\tau_\eta = (\nu/\epsilon)^{1/2}$: Kolmogorov dissipative time; $Re_{\lambda} = (u_0\lambda)/\nu$: Reynolds number based on the Taylor micro-scale; $\lambda = (15\nu u_0^2/\epsilon)^{1/2}$: Taylor micro-scale; $Ro = (\epsilon_f k_f)^{1/3}/\Omega$: Rossby number defined in terms of the energy injection properties, where $k_f = 5$ is the wavenumber where the forcing is acting; f_0 : intensity of the Ornstein-Uhlenbeck forcing; τ_f : decorrelation time of the forcing; $T_0 = u_0/L_0$: Eulerian large-scale eddy turn over time; α : coefficient of the damping term $\alpha\Delta^{-1}u$.

MAX RESOLUTION

$$\partial_t \mathbf{v} + \mathbf{v} \cdot \nabla \mathbf{v} + \mathbf{\Omega} \times \mathbf{v} = \nabla P + \nu \nabla^2 \mathbf{v} + \mathbf{F} - \alpha \mathbf{v}$$

DIMENSIONAL PHENOMENOLOGY

 $k > k_{\Omega} \longrightarrow E(k) \sim \varepsilon^{2/3} k^{-5/3}$ $k_f < k < k_{\Omega} \longrightarrow E(k) \sim (\varepsilon \Omega)^{1/2} k^{-2} \leftrightarrow \overline{\tau_{tr}(k) \sim \frac{\tau_{nl}(k)^2}{\tau_{\Omega}}}$ $k < k_f \longrightarrow E(k) \sim \Omega^2 k^{-3} \leftrightarrow E(k) \sim k^{-5/3}$

