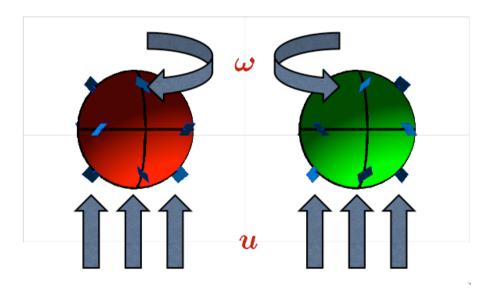
### **Isotropic Helicoids in Complex Flows**

Luca Biferale Dept. of Physics and INFN University of Rome "Tor Vergata" <u>biferale@roma2.infn.it</u>

CREDITS: K. GUSTAFSSON (U. GOTHEBORG ), R. SCATAMACCHIA, F. BONACCORSO (U. TOR VERGATA)





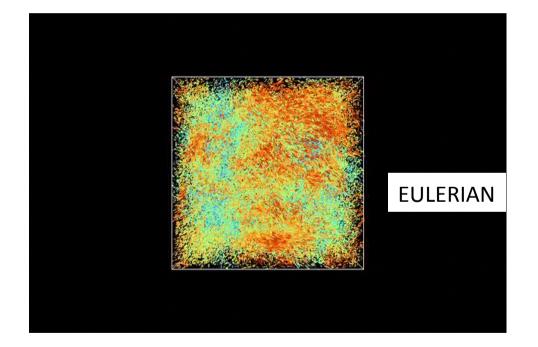


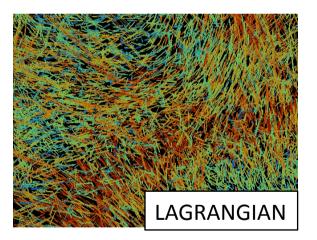




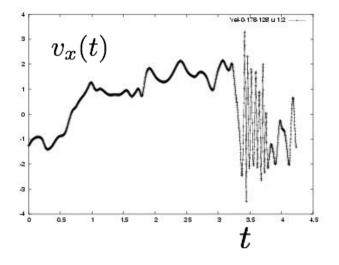


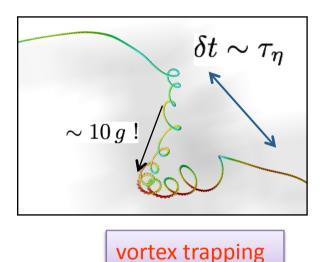
#### **COMPLEX PARTICLES IN COMPLEX FLOWS**: HOW TO ESCAPE/FALL FROM/ON EULERIAN TURBULENT TRAPS?





#### Intermittency - Lagrangian





PHYSICS OF FLUIDS 17, 021701 (2005)

#### Particle trapping in three-dimensional fully developed turbulence

#### L. Biferale

Dipartimento di Fisica and INFN, Università degli Studi di Roma "Tor Vergata," Via della Ricerca Scientifica 1, 00133 Roma, Italy

#### G. Boffetta

Dipartimento di Fisica Generale and INFN, Università degli Studi di Torino, Via Pietro Giuria 1, 10125 Torino, Italy

#### A. Celani CNRS, INLN, 1361 Route des Lucioles, 06560 Valbonne, France

A. Lanotte CNR-ISAC, Str. Prov. Lecce-Monteroni km. 1200, 73100 Lecce, Italy

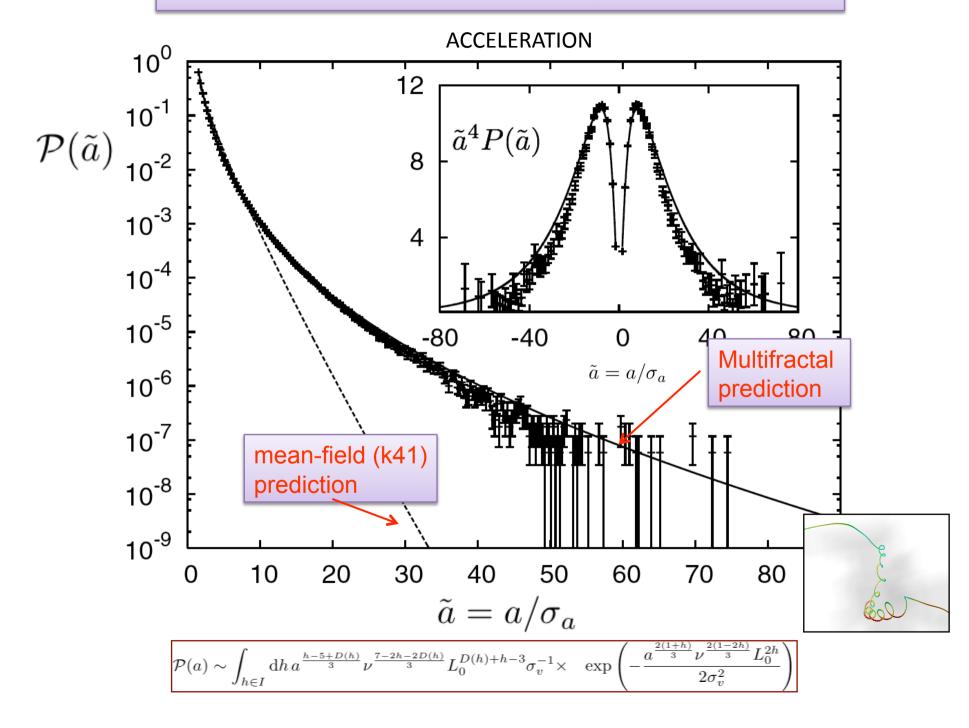
F. Toschi Istituto per le Applicazioni del Calcolo, CNR, Viale del Policlinico 137, 00161 Roma, Italy

#### Lagrangian Properties of Particles in Turbulence

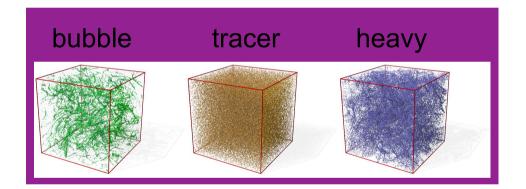
#### Federico Toschi1 and Eberhard Bodenschatz2

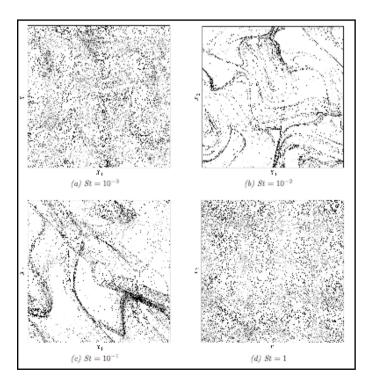
<sup>1</sup>Istituto per le Applicazioni del Calcolo, CNR, I-00161 Rome, Italy; INFN, Sezione di Ferrara, I-44100 Ferrara, Italy; Department of Physics and Department of Mathematics and Computer Science, Eindhoven University of Technology, 5600 MB Eindhoven, The Netherlands; and International Collaboration for Turbulence Research; email: toschi@iac.cnr.it

<sup>2</sup> Max Planck Institute for Dynamics and Self-Organization, D-37077 Goettingen, Germany; Laboratory of Atomic and Solid-State Physics and Sibley School of Mechanical and Aerospace Engineering, Cornell University, Ithaca, New York 14853; Institute for Nonlinear Dynamics, University of Goettingen, D-37073 Goettingen, Germany; and International Collaboration for Turbulence Research

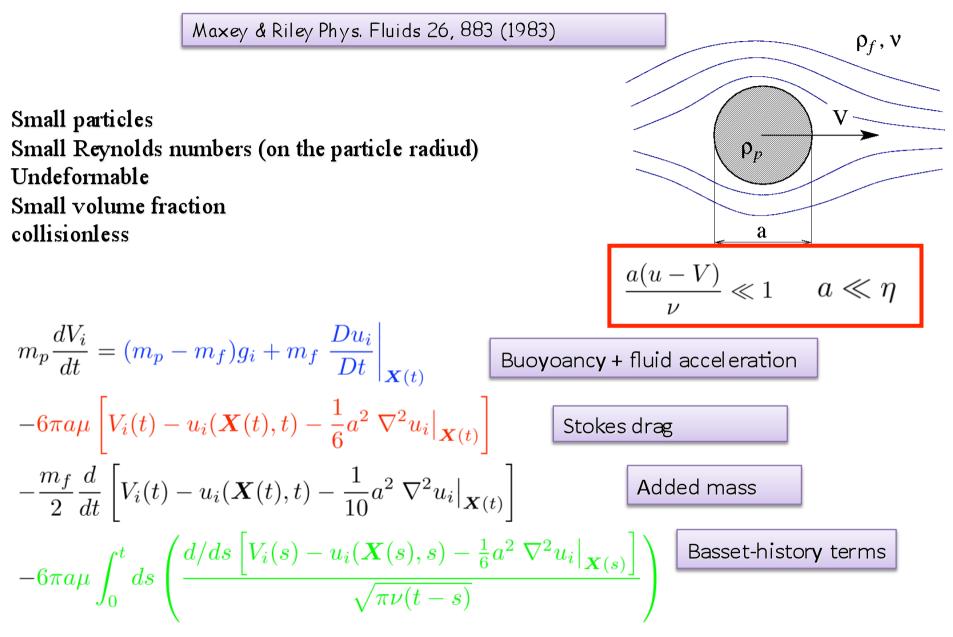


-LAGRANGIAN TURBULENCE -INERTIAL PARTICLES -EFFECTS OF PREFERENTIAL CONCENTRANTIONS -EFFECTS OF CAUSTICS -COMPLEX PARTICLES (TODAY) -SMART PARTICLES

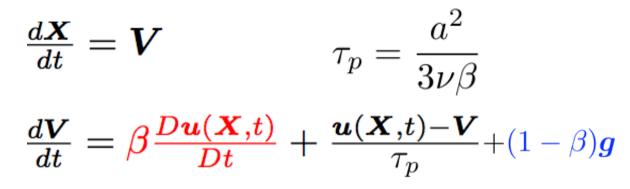




### Eqs of motion for a single particle



### Simplified limit



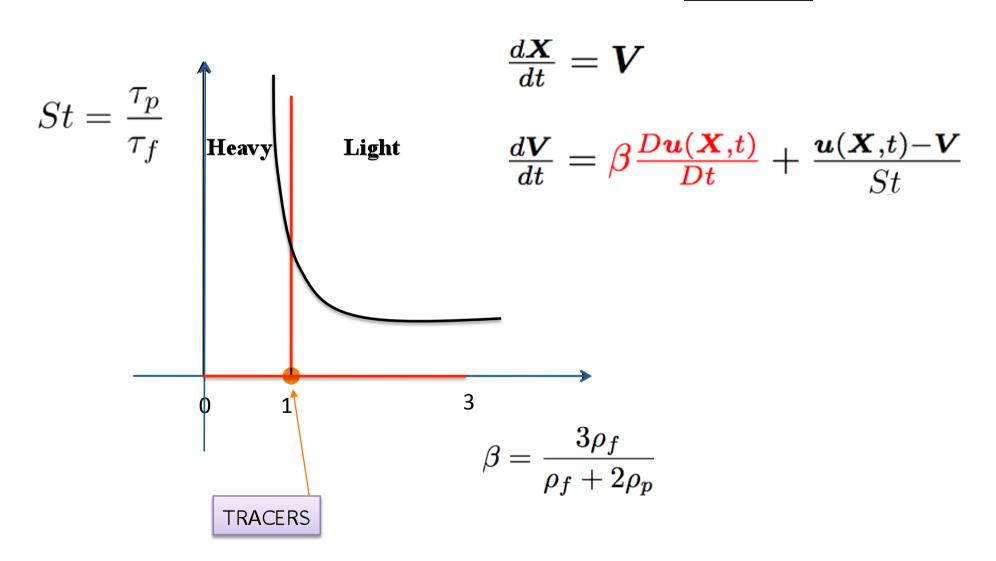
Three-parameters problem

- $au_f$  Fluid characteristic time
- $au_p$  .
  - Particle's characteristic time

$$\begin{array}{ll} \rho_p \gg \rho_f \rightarrow \beta = 0 & \text{Heavy} \\ \rho_f = \rho_p \rightarrow \beta = 1 & \text{tracers} \\ \rho_f \gg \rho_p \rightarrow \beta = 3 & \text{Light} \end{array}$$

**Stokes number** St =Density contrast  $\beta =$ Reynolds Re =





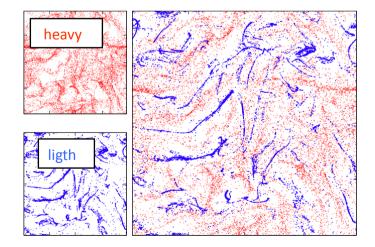
$$\frac{d\mathbf{X}}{dt} = \mathbf{V}$$

$$\frac{d\mathbf{V}}{dt} = \beta \frac{D\mathbf{u}(\mathbf{X},t)}{Dt} + \frac{\mathbf{u}(\mathbf{X},t) - \mathbf{V}}{St}$$

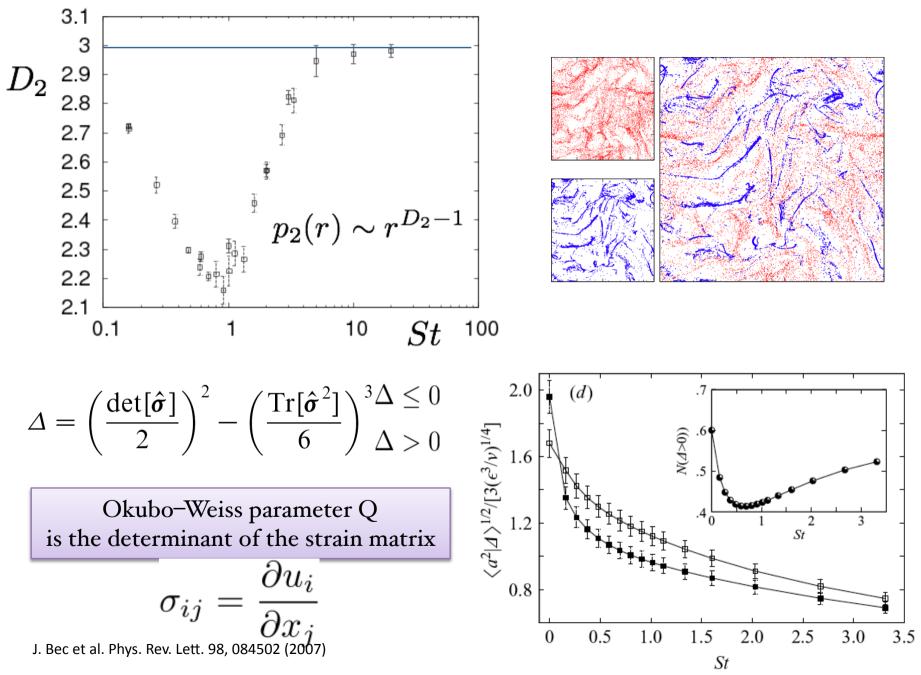
$$V(\boldsymbol{x},t) \approx \boldsymbol{u}(\boldsymbol{x},t) + St(\beta-1)[\partial_t \boldsymbol{u}(\boldsymbol{x},t) + \boldsymbol{u} \cdot \boldsymbol{\nabla} \boldsymbol{u}]$$

$$\boldsymbol{\nabla} \cdot \boldsymbol{V}(\boldsymbol{x}, t) = St(\beta - 1)\boldsymbol{\nabla} \cdot [\boldsymbol{u} \cdot \boldsymbol{\nabla} \boldsymbol{u}] = St(\beta - 1)\sum_{ij} \left(\frac{\partial u_j}{\partial x_i}\right) \left(\frac{\partial u_i}{\partial x_j}\right)$$

$$\begin{split} \beta < 1 \quad S^2 > \Omega^2 \Longrightarrow \boldsymbol{\nabla} \cdot \boldsymbol{V} < 0 \\ \beta > 1 \quad \Omega^2 > S^2 \Longrightarrow \boldsymbol{\nabla} \cdot \boldsymbol{V} < 0 \end{split}$$



### **Preferential Concentration**

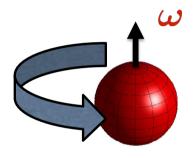




## Spherical particle

Equations for velocity v and angular velocity  $\omega$  for small spherical particle at position r: Happel & Brenner, Low Reynolds number hydrodynamics (1963)

$$\dot{oldsymbol{v}} = rac{1}{ au_{
m p}} \left[ oldsymbol{u}(oldsymbol{r},t) - oldsymbol{v} 
ight] \ \dot{oldsymbol{\omega}} = rac{1}{ au_{
m p}} \left[ rac{10}{3} (oldsymbol{\Omega}(oldsymbol{r},t) - oldsymbol{\omega}) 
ight]$$

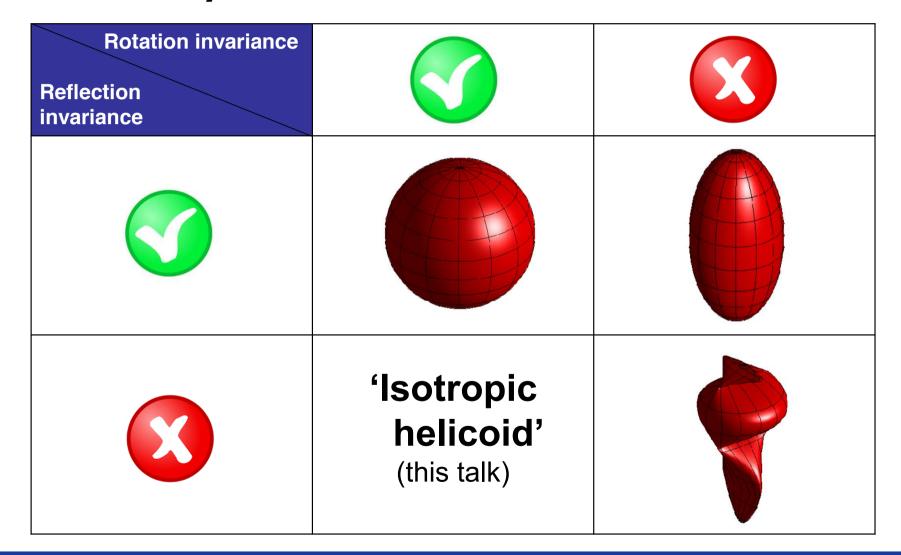


- *u* Fluid velocity
- $\Omega$  Half fluid vorticity
- $au_{\mathrm{p}}$  Particle relaxation time

Dynamics statistically invariant under rotations and reflections if u statistically invariant under rotations and reflections



### Particle symmetries





Recipe from Lord Kelvin:

"An isotropic helicoid can be made by attaching projecting vanes to the surface of a globe in proper positions; for instance cutting at 45° each, at the middles of the twelve quadrants of any three great circles dividing the globe into eight quadrantal triangles."

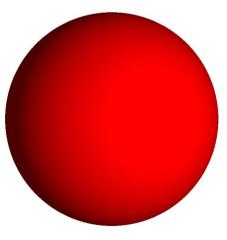
Kelvin, Phil. Mag. **42** (1871)

THES SIMPLEST (BUT NOT SIMPLER) GENERALISATION OF SPHERICAL HEAVY PARTCILES



Recipe from Lord Kelvin (1884)

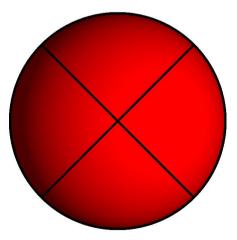
Start with a sphere





Recipe from Lord Kelvin (1884)

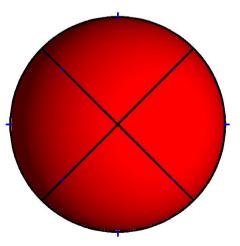
✓ Start with a sphereDraw 3 great circles





Recipe from Lord Kelvin (1884)

✓ Start with a sphere
 ✓ Draw 3 great circles
 Identify 12 vane positions at midpoints of quarter-arcs



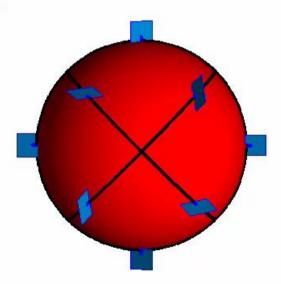
#### UNIVERSITA' dogii studi di Roma tor vercata

# Example of an isotropic helicoid

Recipe from Lord Kelvin (1884)

- $\checkmark$  Start with a sphere
- ✓ Draw 3 great circles
- $\checkmark$  Identify 12 vane positions at midpoints of quarter-arcs

Put a vane on each vane position (45° to arc line)

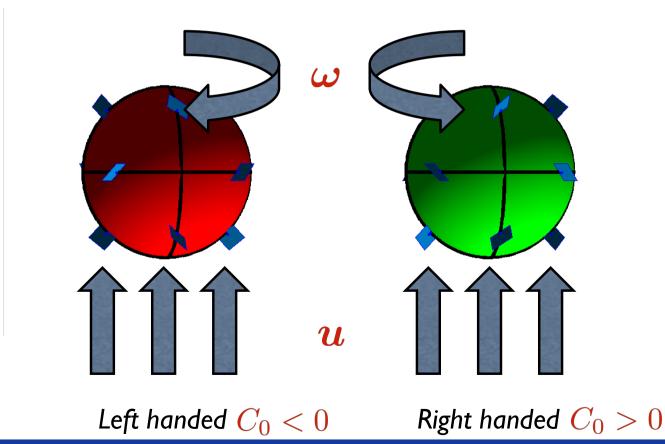




# Chirality

In a constant flow u, the isotropic helicoid starts spinning around the flow direction with angular velocity  $\omega$ .

The spinning direction depends on the chirality of the vanes.



# Motion of an 'isotropic helicoid'

Equations for velocity v and angular velocity  $\omega$  for small isotropic helicoid: Happel & Brenner, Low Reynolds number hydrodynamics (1963)

$$\dot{\boldsymbol{v}} = \frac{1}{\tau_{\rm p}} \left[ \boldsymbol{u}(\boldsymbol{r}, t) - \boldsymbol{v} + \frac{2a}{9} C_0(\boldsymbol{\Omega}(\boldsymbol{r}, t) - \boldsymbol{\omega}) \right]$$
$$\dot{\boldsymbol{\omega}} = \frac{1}{\tau_{\rm p}} \left[ \frac{10}{3} (\boldsymbol{\Omega}(\boldsymbol{r}, t) - \boldsymbol{\omega}) + \frac{5}{9a} C_0(\boldsymbol{u}(\boldsymbol{r}, t) - \boldsymbol{v}) \right]$$

Stokes' law translation – rotation coupling (scalar)

 $a = \sqrt{5I_0/(2m)}$  Particle 'size' (defined by mass *m* and moment of inertia  $I_0$ )

 $C_0$  Helicoidality

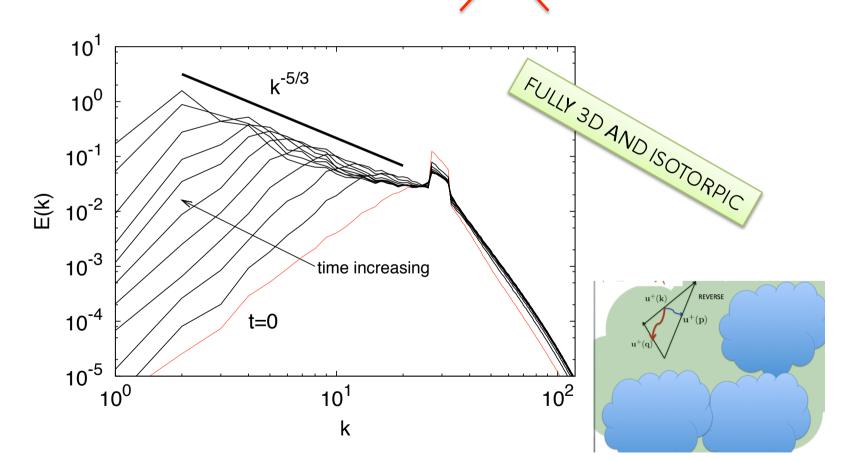
RSITA' degil Studi di Roma

Ratio of rotational and translational inertia fixed to that of sphere

Equations break spatial reflection symmetry ( $\omega$  pseudovector)

REGIONS WITH STRONGLY MULTI-SCALE HELICAL CHOERENCY -> REVERT ENERGY CASCADE

$$\begin{cases} E = \sum_{k} |u^{+}(k)|^{2} + |u^{-}(k)|^{2}; \\ H = \sum_{k} k(|u^{+}(k)|^{2} - |u^{-}(k)|^{2}). \end{cases}$$



L.B., S. MUSACCHIO & F. TOSCHI Phys. Rev. Lett. 108 164501, 2012.

## **Dimensionless** parameters

Stokes number  $\operatorname{St} \equiv \frac{\tau_{\mathrm{p}}}{\tau_{\eta}}$  Size  $\overline{a} \equiv \frac{a}{\eta}$  Helicoidality  $C_0$ 

with  $au_{\eta}$  and  $\eta$  smallest time- and length scales of flow.

Dynamics may grow indefinitely unless  $-\sqrt{27} < C_0 < \sqrt{27}$  .

St and  $\overline{a}$  constrained by particle density higher than that of the fluid and geometrical size must be smaller than  $\eta$ .

Simulations and theory is done using a random single-scale flow characterised by the Kubo number

$$\mathrm{Ku} \equiv \frac{u_0 \tau_{\eta}}{\eta}$$
 with  $u_0$  typical speed of flow.

#### UNIVERSITA' degli STUDI di ROMA T O R V E R G A T A

# Clustering at small St

Expand compressibility of particle-velocity field  $oldsymbol{
abla}\cdotoldsymbol{v}$  in small  $\mathrm{St}\sim au_\mathrm{p}$ 

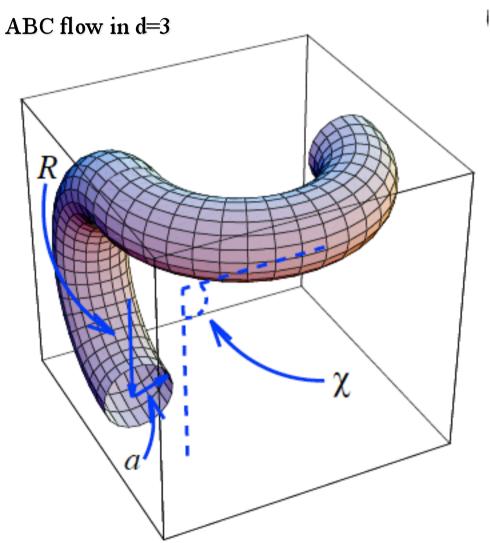
$$\boldsymbol{\nabla} \cdot \boldsymbol{v} = -\frac{27}{27 - C_0^2} \tau_{\mathrm{p}} \left[ \mathrm{Tr} (\boldsymbol{\nabla} \boldsymbol{u}^{\mathrm{T}} \boldsymbol{\nabla} \boldsymbol{u}^{\mathrm{T}}) - \frac{1}{15} \mathrm{aC}_0 \mathrm{Tr} (\boldsymbol{\nabla} \boldsymbol{u}^{\mathrm{T}} \boldsymbol{\nabla} \boldsymbol{\Omega}^{\mathrm{T}}) \right]$$

Centrifuge effect with modified amplitude Maxey, J. Fluid Mech. **174** (1987) Term due to parity breaking of system

Reflection-invariant systems have  $\langle \operatorname{Tr}(\nabla u^{\mathrm{T}} \nabla \Omega^{\mathrm{T}}) \rangle = 0$ Isotropic helicoids violate that relation  $\langle \operatorname{Tr}(\nabla u^{\mathrm{T}} \nabla \Omega^{\mathrm{T}}) \rangle \propto \tau_{\mathrm{p}} C_{0}$ 

 $\Rightarrow$  In a parity-invariant isotropic flow clustering does not depend on sign of  $C_0$ 

#### Eulerian smooth but Lagrangian non-trivial



 $\dot{x} = A\sin z + C\cos y,$ 

$$\dot{y} = B\sin x + A\cos z,$$

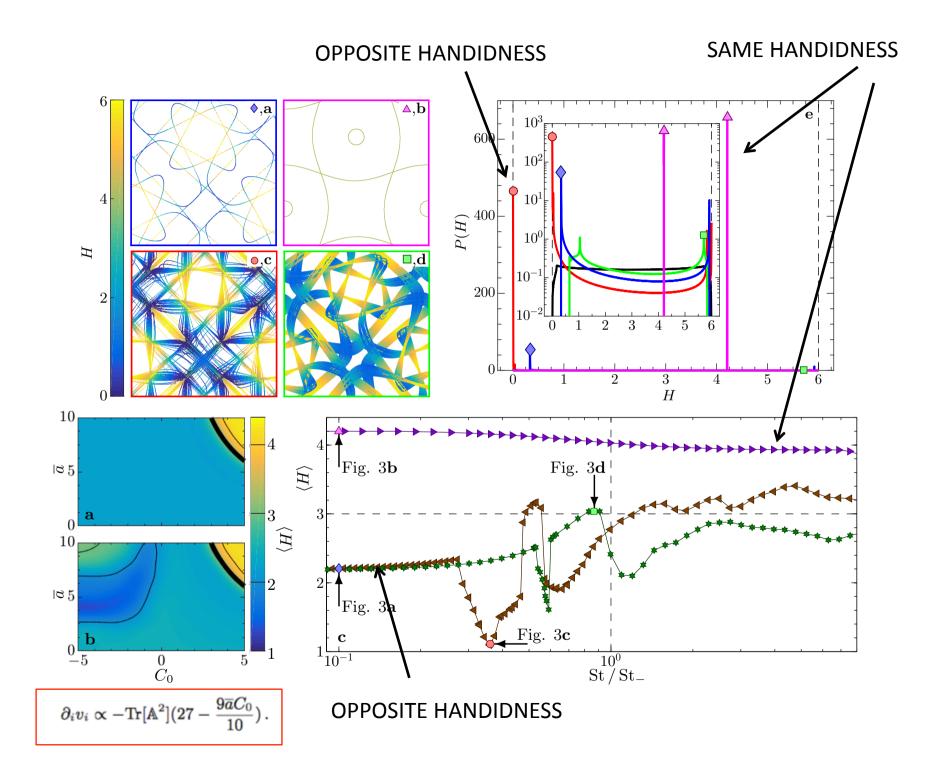
$$\dot{z} = C\sin y + B\cos x$$

$$\mathbf{v} \parallel \omega$$

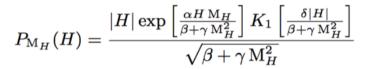
Exact stationaty solution of Euler equation

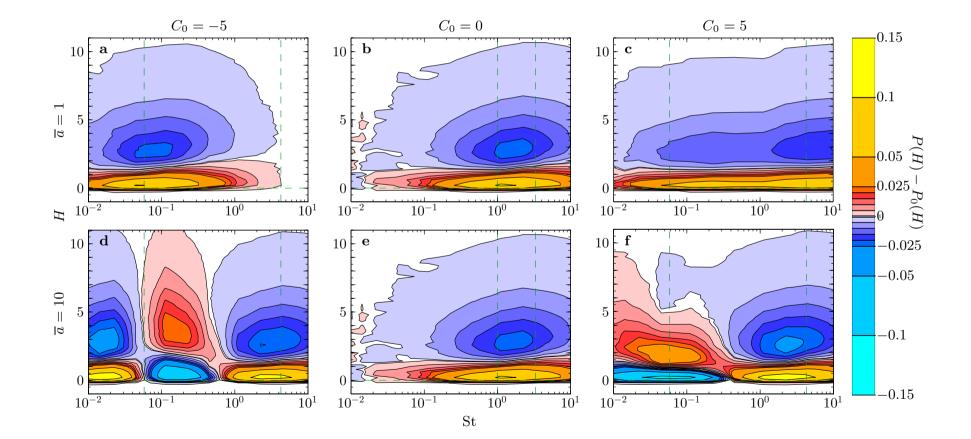
$$\partial_i v_i \propto - {
m Tr}[{\mathbb A}^2] (27 - {9 \overline{a} C_0 \over 10}) \, .$$

**HELICOIDS MIGHT BEHAVE AS LIGHT OR HEAVY PARTICLES !!!** 



#### STOCHASTIC HELICAL FLOW





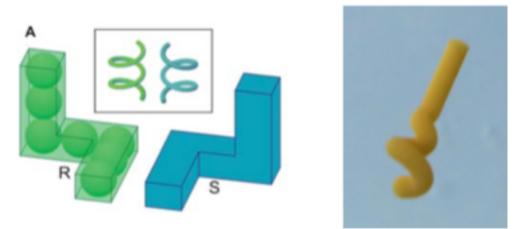


Figure 2. (Lafe) Example of skiral particles that have been used in microfluidic experiments. Such particles early

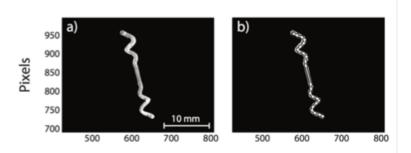
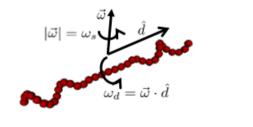


Figure 5: An example of a complex particle, a "strain probe" is basically a chiral-dipole and is sensitive to the local strain level in turbulence. This type of 3D printed particles have been designed and tracked for both position and orientation in turbulent flows by means of optical techniques [20], [24]. Similarly shaped particles were studied numerically by means of Stokesian dynamics simulations [25] (see Figure 6).





## Conclusions

Isotropic helicoids are rotation invariant particles which break reflection invariance (two chiralities)

Coupling between translational and rotational degrees of freedom changes dynamics compared to spherical particles (modified clustering, preferential sampling etc.)

The two chiralities may show different dynamics if the particle size is not too small and flow is persistent

Flows with broken parity invariance increase the differences in the dynamics of the two particles

K. GUSTAFSSON AND L.B. PHYS. REV. FLUIDS (IN PRESS) 2016.